

GRAIN GROWTH, OPTICAL POLARISATION AND EXTINCTION IN INTERSTELLAR CLOUDS

A P Jones Department of Mathematics, UMIST, PO Box 88, Manchester M60 1QD, England, UK

Increases in the wavelength of maximum polarisation and the ratio of total to selective extinction are generally assumed to be the result of grain growth in interstellar clouds (McMillan, 1978; Whittet and Blades, 1980; Clayton and Mathis, 1988). Using a grain model in which the interstellar extinction is explained by amorphous carbon (a-C) and hydrogenated amorphous carbon (a-C:H) mantles on silicate cores (Duley, 1987; Jones, Duley and Williams, 1987; Duley, Jones and Williams, 1988) we have theoretically studied the effects of these absorbing carbonaceous mantles on the polarising properties of large (500 Å - 2500 Å radius) silicate grains. We adopt the polarisation model of Mathis (1986) and show that carbon-coated silicate grains can explain the relationship between the wavelength of maximum polarisation (λ_{max}) and R observed toward dark clouds.

References

Clayton, G.C. and Mathis, J S.: 1988, Ap. J. 327, 911.

Duley, W.W.: 1987, Mon. Not. R. astr. Soc. 229, 203.

Duley, W.W., Jones, A.P. and Williams, D.A.: 1987, Mon. Not. R. astr. Soc. 229, 213.

Jones, A.P., Duley, W.W. and Williams, D.A.: 1988, Mon. Not. R. astr. Soc. in press.

McMillan, R.S.:1978, Ap. J. 225, 880.

Whittet, D. C. B. and Blades, J. C.: 1980, Mon. Not. R. astr. Soc. 191, 309.